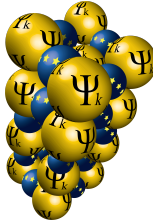


# ICTP/Psi-k/CECAM School on Electron-Phonon Physics from First Principles

Trieste, 19-23 March 2018



Lecture Fri.3

# Superconducting gap and critical temperature using EPW

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# Lecture Summary

- Structure of the code
- Technicalities and convergences parameters

# Migdal-Eliashberg Equations on the Imaginary Axis

$$Z_{n\mathbf{k}}(i\omega_j) = 1 + \frac{\pi T}{\omega_j N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\omega_{j'}}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}}$$

mass renormalization function  $\times \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F)$

$$Z_{n\mathbf{k}}(i\omega_j) \Delta_{n\mathbf{k}}(i\omega_j) = \frac{\pi T}{N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\Delta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}}$$

superconducting gap function  $\times [\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) - \mu_c^*] \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F)$

↑  
anisotropic e-ph coupling strength

# Migdal-Eliashberg Equations on the Imaginary Axis

Input variables:

```
1 ephwrite = .true.
```

CALL write\_ephmat(...)

$$\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j) = N_F \sum_{\nu} \int_0^{\infty} d\omega \frac{2\omega}{\omega_j^2 + \omega^2} |g_{m\nu}(\mathbf{k}, \mathbf{q})|^2 \delta(\omega - \omega_{\mathbf{q}\nu})$$

↑  
e-ph matrix elements  
on fine  $\mathbf{k}$  and  $\mathbf{q}$  grids  
(prefix.ephmatX files)

# Structure of the Code

The epw.f90 file:

```
1 CALL elphon_shuffle_wrap()  
2 --> CALL ephwann_shuffle ( nqc , xqc )
```

The ephwann\_shuffle.f90 file:

```
1 IF (ephwrite) THEN  
2   IF ( iq .eq. 1 ) THEN  
3     ! Find (irreducible) k-points on the fine grid within  
4       the Fermi window and write prefix.ikmap file  
5     CALL kmesh_fine  
6     ! Compute the index of k+q on the fine k-grid  
7     CALL kqmap_fine  
8   ENDIF  
9   ! Write prefix.ephmatX, prefix.freq, and prefix.egnv files  
10  CALL write_ephmat( iq )  
11 ENDIF
```

# Structure of the Code

Input variables:

```
1 ephwrite = .true.          fsthick = 0.4 ! eV
2                               degaussw = 0.1 ! eV
3 mp_mesh_k = .true. ! use irreducible k-points
4 nkf1 = 20                   nqf1 = 20
5 nkf2 = 20                   nqf2 = 20
6 nkf3 = 20                   nqf3 = 20
```

The **fine** **k** and **q** grids are required to be **uniform** and **commensurate** such that the  $\mathbf{k}' = \mathbf{k} + \mathbf{q}$  grid maps into the **k** grid.

Files created (used for solving the Migdal-Eliashberg equations):

```
1 prefix.ephmatX      ! e-ph matrix elements within the Fermi
   window for the fine k and q grids (X=#files=#processors)
2 prefix.freq         ! Phonon frequencies on the q fine grid
3 prefix.egn         ! Eigenvalues within the Fermi window on
   the fine k grid
4 prefix.ikmap        ! Index of each k-point on the uniform grid
   on the corresponding irreducible grid
```

# Electron-phonon Coupling Strength

Input variables:

```
1 eliashberg = .true.
```

CALL lambdar\_aniso\_ver1(...)

$$\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j) = N_F \sum_{\nu} \int_0^{\infty} d\omega \frac{2\omega}{\omega_j^2 + \omega^2} |g_{m\nu}(\mathbf{k}, \mathbf{q})|^2 \delta(\omega - \omega_{\mathbf{q}\nu})$$

$$\lambda_{n\mathbf{k}}(\omega_j) = \sum_m \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F)}{N_F} \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j)$$

$$\lambda(\omega_j) = \sum_n \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \frac{\delta(\epsilon_{n\mathbf{k}} - \epsilon_F)}{N_F} \lambda_{n\mathbf{k}}(\omega_j)$$



# Structure of the Code

The epw.f90 file:

```
1 IF ( eliashberg ) THEN
2     CALL eliashberg_eqs()
3 ENDIF
```

The eliashberg.f90 file:

```
1 IF ( .not. liso .AND. .not. laniso ) THEN
2     CALL read_frequencies ! Read prefix.freq
3     CALL read_eigenvalues ! Read prefix.egnv
4     CALL read_kqmap       ! Read prefix.ikmap and compute
                           ! the index of k+q on the fine k-grid
5     CALL read_ephmat      ! Read prefix.ephmatX
6     !
7     CALL eliashberg_init
8     CALL evaluate_a2f_lambda
9     --> CALL lambdar_aniso_ver1(...)
10    CALL estimate_tc_gap
11 ENDIF
```

# Structure of the Code

## Files created:

```
1 prefix.a2f ! Eliashberg spectral function as a function of
    frequency (meV) for various smearings
2 prefix.a2f_iso ! 2nd column is the Eliashberg spectral
    function corresponding to the first smearing in .a2f.
    Remaining columns are the mode-resolved Eliashberg
    spectral function (there is no specific information on
    which modes correspond to which atomic species).
3 prefix.lambda_k_pairs ! \lambda_nk distribution on FS
4 prefix.lambda_FS ! k-point Cartesian coords, n, E_nk-E_F[eV],
    \lambda_nk
5 prefix.phdos ! Phonon DOS (same as .a2f)
6 prefix.phdos_proj ! Phonon DOS (same as .a2f_iso)
```

# Structure of the Code

## Files created:

```
1 prefix.a2f ! Eliashberg spectral function as a function of
    frequency (meV) for various smearings
2 prefix.a2f_iso ! 2nd column is the Eliashberg spectral
    function corresponding to the first smearing in .a2f.
    Remaining columns are the mode-resolved Eliashberg
    spectral function (there is no specific information on
    which modes correspond to which atomic species).
3 prefix.lambda_k_pairs ! \lambda_nk distribution on FS
4 prefix.lambda_FS ! k-point Cartesian coords, n, E_nk-E_F[eV],
    \lambda_nk
5 prefix.phdos ! Phonon DOS (same as .a2f)
6 prefix.phdos_proj ! Phonon DOS (same as .a2f_iso)
```

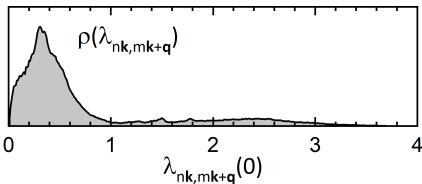
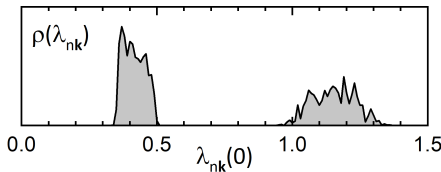
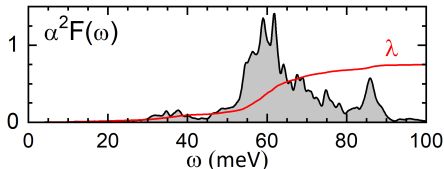
## Files created with iverbosity = 2:

```
1 prefix.lambda_aniso ! E_nk-E_F[eV], \lambda_nk, k, n
2 prefix.lambda_pairs ! \lambda_nk,mk+q distribution on FS
3 prefix.lambda_YY.cube ! Same as prefix.lambda_FS for VESTA
    visualization. YY is the band index within the energy
    window
```

# Electron-Phonon Coupling Strength

Files created (eliashberg = .true.):

```
1 prefix.a2f
2 prefix.lambda_k_pairs
3 prefix.lambda_pairs ! (iverbosity = 2)
```



Figures adapted from Margine and F. Giustino, Phys. Rev. B 87, 024505 (2013)

# Migdal-Eliashberg Equations on the Imaginary Axis

Input variables:

```
1 eliashberg      = .true.
2 laniso          = .true.
3 limag           = .true.
```

CALL eliashberg\_aniso\_iaxis

$$\begin{aligned} Z_{n\mathbf{k}}(i\omega_j) &= 1 + \frac{\pi T}{\omega_j N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\omega_{j'}}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}} \\ \text{mass renormalization} & \\ \text{function} & \times \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F) \end{aligned}$$

$$\begin{aligned} Z_{n\mathbf{k}}(i\omega_j) \Delta_{n\mathbf{k}}(i\omega_j) &= \frac{\pi T}{N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\Delta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}} \\ \text{superconducting} & \\ \text{gap function} & \times [\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) - \mu_c^*] \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F) \end{aligned}$$

CALL lambdar\_aniso\_ver1(...)

# Structure of the Code

The epw.f90 file:

```
1 IF ( eliashberg ) THEN
2     CALL eliashberg_eqs()
3 ENDIF
```

The eliashberg.f90 file:

```
1 IF ( laniso ) THEN
2     CALL read_frequencies ! Read prefix.freq
3     CALL read_eigenvalues ! Read prefix.egnv
4     CALL read_kqmap       ! Read prefix.ikmap and compute
                           ! the index of k+q on the fine k-grid
5     CALL read_ephmat      ! Read prefix.ephmatX
6
7     CALL eliashberg_init
8     CALL evaluate_a2f_lambda
9     CALL estimate_tc_gap
10    IF ( limag ) CALL eliashberg_aniso_iaxis
11 ENDIF
```

# Structure of the Code

The eliashberg\_aniso\_iaxis.f90 file:

```
1 DO itemp = 1, nstemp ! loop over temperature
2   ! Generate the frequency grid on the imaginary axis
3   CALL gen_freqgrid_iaxis( itemp )
4   IF ( ( limag .AND. .not. imag_read ) .OR. ( limag .AND.
5     imag_read .AND. itemp .ne. 1 ) ) THEN
6     iter = 1
7     conv = .false.
8     DO WHILE ( .not. conv .AND. iter .le. nsiter )
9       ! Solve Migdal-Eliashberg eqs on the imaginary axis
10      CALL sum_eliashberg_aniso_iaxis( itemp, iter, conv )
11      --> CALL kernel_aniso_iaxis( itemp )
12      --> CALL lambdar_aniso_ver1 ( ... )
13      iter = iter + 1
14    ENDDO ! iter
15    IF ( conv ) CALL free_energy( itemp )
16  ELSEIF ( limag .AND. imag_read .AND. itemp .eq. 1 ) THEN
17    ! Read from file Delta and Znrm on the imaginary axis
18    CALL eliashberg_read_aniso_iaxis( itemp )
19  ENDIF
20 ENDDO
```

# Structure of the Code

The eliashberg\_aniso\_iaxis.f90 file:

```
1 DO itemp = 1, nstep ! loop over temperature
2   ! Generate the frequency grid on the imaginary axis
3   CALL gen_freqgrid_iaxis( itemp )
4   IF ( ( limag .AND. .not. imag_read ) .OR. ( limag .AND.
5     imag_read .AND. itemp .ne. 1 ) ) THEN
6     ! Solve Migdal-Eliashberg eqs on the imaginary axis
7     ....
8   ELSEIF ( limag .AND. imag_read .AND. itemp .eq. 1 ) THEN
9     ! Read from file Delta and Znorm on the imaginary axis
10    CALL eliashberg_read_aniso_iaxis( itemp ) <--
11  ENDIF
12 ENDDO
```

Restart option:

- `imag_read = .true.`



# Structure of the Code

## Input variables:

```
1 eliashberg = .true.          conv_thr_iaxis = 1.0d-4
2 limag      = .true.          nsiter = 500
3 laniso     = .true.          wscut  = 1.0 ! eV cutoff freq.
4
5 nstemp     = 4                muc    = 0.16 ! Coulomb parameter
6 tempmin   = 15.0
7 tempmax   = 60.0
```

# Structure of the Code

Input variables:

```
1 eliashberg = .true.          conv_thr_iaxis = 1.0d-4
2 limag      = .true.          nsiter = 500
3 laniso     = .true.          wscut  = 1.0 ! eV cutoff freq.
4
5 nstemp     = 4                muc     = 0.16 ! Coulomb parameter
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7 tempsmax   = 60.0
```

Input variables:

```
1 eliashberg = .true.          conv_thr_iaxis = 1.0d-4
2 limag      = .true.          nsiter = 500
3 laniso     = .true.          wscut  = 1.0 ! eV cutoff freq.
4
5 temps(1)   = 15.0            muc     = 0.16 ! Coulomb parameter
6 temps(2)   = 30.0
7 temps(3)   = 45.0
8 temps(4)   = 60.0
```

# Structure of the Code

Files created (XX indicates the temperature):

```
1 prefix.imag_aniso_XX           ! w_j[eV], E_nk-E_F[eV], Z_nk,
   \Delta_nk[eV], Z^N_nk
2 prefix.imag_aniso_gap0_XX      ! \Delta_nk(0)[meV]
   distribution on FS
3 prefix.imag_aniso_gap_FS_XX    ! k-point Cartesian coords,
   band index within energy window, E_nk-E_F[eV],
   \Delta_nk(0)[eV]
```

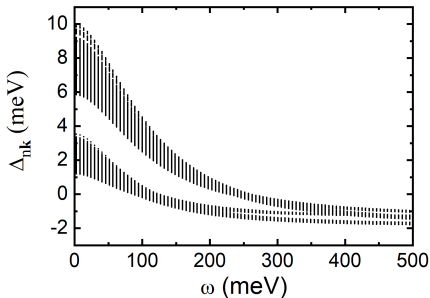
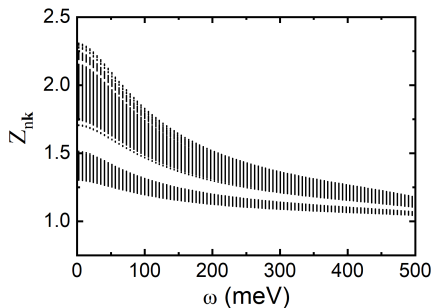
Files created with iverbosity = 2 (YY is the band index within the energy window):

```
1 prefix.imag_aniso_gap0_XX_YY.cube ! Same as prefix.
   imag_aniso_gap_FS_XX for VESTA visualization
```

# Migdal-Eliashberg Equations on the Imaginary Axis

Files created:

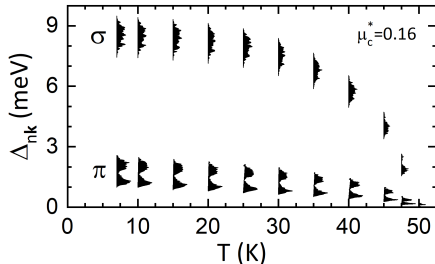
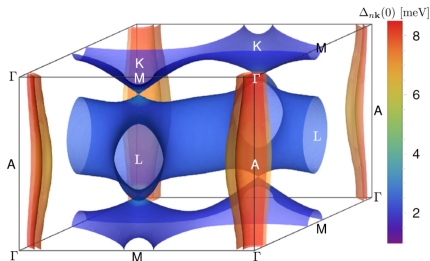
```
1 prefix.imag_aniso_XX
```



# Superconducting Gap and Critical Temperature

Files created:

```
1 prefix.imag_aniso_gap0_XX_YY.cube ! uniform k-grid,  
  \Delta_nk(0) [eV]  
2 prefix.imag_aniso_gap0_XX      ! \Delta_nk(0) [eV]  
  distribution on FS
```



Left and right figures from Poncé et al, *Comp. Phys. Commun.* 209, 116 (2016) and Margine and Giustino, *Phys. Rev. B* 87, 024505 (2013)

# Superconducting Gap Convergence

Convergence must be made for:

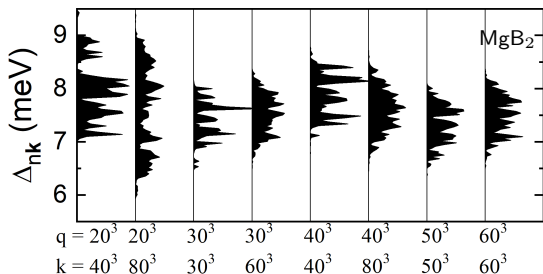
1	<code>nkf1, nkf2, nkf3</code>	<code>wscut</code>
2	<code>nqf1, nqf2, nqf3</code>	<code>fsthick</code>

Left figure from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

# Superconducting Gap Convergence

Convergence must be made for:

1	nkf1, nkf2, nkf3	wscut
2	nqf1, nqf2, nqf3	fsthick

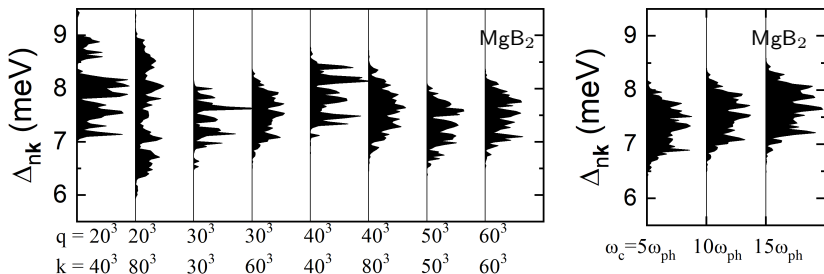


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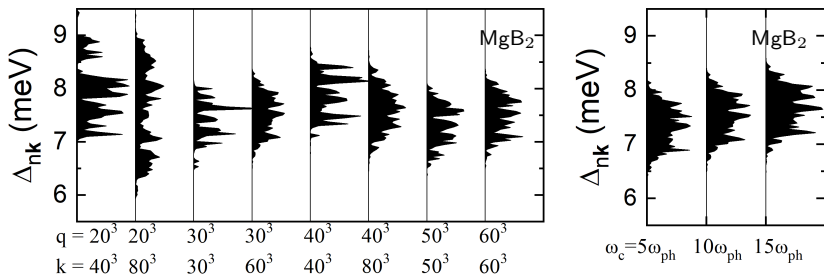
Left figure from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)



# Superconducting Gap Convergence

Convergence must be made for:

1	nkf1, nkf2, nkf3	wscut
2	nqf1, nqf2, nqf3	fsthick



Description of anisotropic quantities requires very dense k and q grids

Left figure from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

# Electron-Phonon Coupling Strength Convergence

Reference	$\mathbf{q}$ -mesh	$\mathbf{k}$ -mesh	$\lambda$
Bohnen <i>et al.</i> (Ref. 73)	$6^3$	$36^3$	0.73
Choi <i>et al.</i> (Ref. 17)	$12 \times 18^2$	$12 \times 18^2$	0.73
Floris <i>et al.</i> (Ref. 15)	$20^3$	$24^3$	0.71
Eiguren <i>et al.</i> (Ref. 74)	$40^3$	$40^3$	0.776
Calandra <i>et al.</i> (Ref. 75)	$20^3$	$80^3$	0.741
This work	$20^3$	$40^3$	0.735
	$20^3$	$80^3$	0.739
	$30^3$	$30^3$	0.782
	$30^3$	$60^3$	0.748
	$40^3$	$40^3$	0.735
	$40^3$	$80^3$	0.739
	$50^3$	$50^3$	0.744
	$60^3$	$60^3$	0.748

Table from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

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	$30^3$	$60^3$	0.748
	$40^3$	$40^3$	0.735
	$40^3$	$80^3$	0.739
	$50^3$	$50^3$	0.744
	$60^3$	$60^3$	0.748

Isotropic quantities are less sensitive to the size of  $\mathbf{k}$  and  $\mathbf{q}$  grids

Table from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

# Superconducting Specific Heat

- **Free energy difference** between superconducting and normal states

$$\Delta F = -\pi T \sum_{nj} \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \left[ \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2(i\omega_j)} - |\omega_j| \right] \\ \times \left[ Z_{n\mathbf{k}}(i\omega_j) - Z_{n\mathbf{k}}^N(i\omega_j) |\omega_j| / \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2(i\omega_j)} \right] \delta(\epsilon_{n\mathbf{k}} - \epsilon_F)$$

Figure from Poncé et al, Comp. Phys. Commun. 209, 116 (2016)

# Superconducting Specific Heat

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$$\Delta F = -\pi T \sum_{nj} \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \left[ \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2} - |\omega_j| \right] \\ \times \left[ Z_{n\mathbf{k}}(i\omega_j) - Z_{n\mathbf{k}}^N(i\omega_j) |\omega_j| / \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2} \right] \delta(\epsilon_{n\mathbf{k}} - \epsilon_F)$$

$$\Delta C = -T \frac{d^2 \Delta F}{dT^2}$$

↑  
superconducting  
specific heat

Figure from Poncé et al, Comp. Phys. Commun. 209, 116 (2016)

# Superconducting Specific Heat

- Free energy difference between superconducting and normal states

$$\Delta F = -\pi T \sum_{nj} \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \left[ \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2} (i\omega_j) - |\omega_j| \right] \\ \times \left[ Z_{n\mathbf{k}}(i\omega_j) - Z_{n\mathbf{k}}^N(i\omega_j) |\omega_j| / \sqrt{\omega_j^2 + \Delta_{n\mathbf{k}}^2} \right] \delta(\epsilon_{n\mathbf{k}} - \epsilon_F)$$

$$\Delta C = -T \frac{d^2 \Delta F}{dT^2}$$

↑  
superconducting  
specific heat

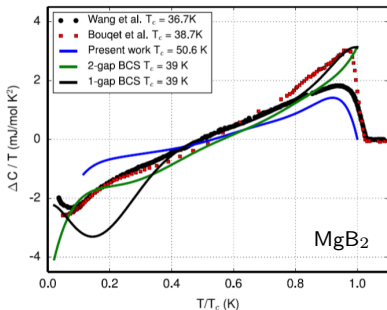


Figure from Poncé et al, Comp. Phys. Commun. 209, 116 (2016)

# Structure of the Code

The eliashberg\_aniso\_iaxis.f90 file:

```
1 DO itemp = 1, nstemp ! loop over temperature
2   CALL gen_freqgrid_iaxis( itemp )
3   IF ( ( limag .AND. .not. imag_read ) .OR. ( limag .AND.
4     imag_read .AND. itemp .ne. 1 ) ) THEN
5     iter = 1
6     conv = .false.
7     DO WHILE ( .not. conv .AND. iter .le. nsiter )
8       CALL sum_eliashberg_aniso_iaxis( itemp, iter, conv )
9       iter = iter + 1
10      ENDDO ! iter
11      IF ( conv ) CALL free_energy( itemp ) <--
12      ELSEIF ( limag .AND. imag_read .AND. itemp .eq. 1 ) THEN
13        CALL eliashberg_read_aniso_iaxis( itemp )
14      ENDDO
```

Files created :

```
1 prefix.fe_XX ! temperature, free energy difference between
   superconducting and normal state
```

## Migdal-Eliashberg Equations on Real Axis

- The Migdal-Eliashberg equations on the imaginary frequency axis are computationally efficient (only involve sums over a finite number of Matsubara frequencies) and they are adequate for calculating the critical temperature and the superconducting gap.



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# Migdal-Eliashberg Equations on Real Axis

- The Migdal-Eliashberg equations on the imaginary frequency axis are computationally efficient (only involve sums over a finite number of Matsubara frequencies) and they are adequate for calculating the critical temperature and the superconducting gap.
- To extract information about the spectral properties (e.g., the quasi particle density of states or the single particle excitation spectrum), we need to solve the Migdal-Eliashberg equations on the real energy axis.
- Direct evaluation of the Migdal-Eliashberg equations on the real energy axis is in principle possible but very demanding computationally since it involves the evaluation of many principal value integrals.

# Migdal-Eliashberg Equations on Real Axis

- The Migdal-Eliashberg equations on the imaginary frequency axis are computationally efficient (only involve sums over a finite number of Matsubara frequencies) and they are adequate for calculating the critical temperature and the superconducting gap.
- To extract information about the spectral properties (e.g., the quasi particle density of states or the single particle excitation spectrum), we need to solve the Migdal-Eliashberg equations on the real energy axis.
- Direct evaluation of the Migdal-Eliashberg equations on the real energy axis is in principle possible but very demanding computationally since it involves the evaluation of many principal value integrals.
- As an alternative, the solutions on the real energy axis can be obtained by analytic continuation of the solutions along the imaginary frequency axis. The analytic continuation can be performed by using Padé approximants (very light computationally) or by means of an iterative procedure (very heavy computationally).

# Structure of the Code

The eliashberg\_aniso\_iaxis.f90 file:

```
1 DO itemp = 1, nstemp ! loop over temperature
2   ...
3   ...
4   IF ( lpade ) THEN
5     conv = .false.
6     CALL pade_cont_aniso_iaxis_to_raxis( itemp, N, conv )
7     IF ( conv ) CALL dos_quasiparticle( itemp )
8   ENDIF
9   IF ( lacon ) THEN
10    iter = 1
11    conv = .false.
12    DO WHILE ( .not. conv .AND. iter .le. nsiter )
13      CALL analytic_cont_aniso_iaxis_to_raxis( itemp, iter
14        , conv )
15      iter = iter + 1
16    ENDDO ! iter
17    IF ( conv ) CALL dos_quasiparticle( itemp )
18  ENDIF
19 ENDDO
```

# Structure of the Code

## Input variables:

```
1 eliashberg = .true.          conv_thr_iaxis = 1.0d-4
2 limag      = .true.          nsiter = 500
3 laniso     = .true.          wscut  = 1.0 ! eV cutoff freq.
4 lpade      = .true.
5 lacon      = .true.          conv_thr_racon = 1.0d-4
6
7 nstep      = 4                muc     = 0.16 ! Coulomb parameter
8 tempsmin   = 15.0
9 tempsmax   = 60.0
```

# Structure of the Code

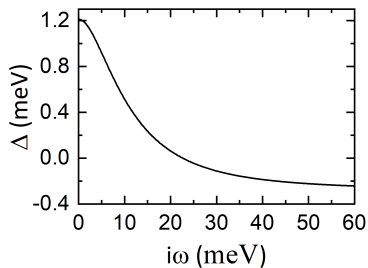
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```

## Files created (XX indicates the temperature):

```
1 prefix.pade_aniso_XX ! w[eV], E_nk-E_F[eV], RE[Z_nk], Im[Z_nk
   ], Re[\Delta_nk][eV], Im[\Delta_nk][eV] (iverbosity=2)
2 prefix.pade_aniso_gap0_XX ! Re[\Delta_nk(0)][eV] distribution
   on FS
3 prefix.acon_aniso_XX ! w[eV], E_nk-E_F[eV], RE[Z_nk], Im[Z_nk
   ], Re[\Delta_nk][eV], Im[\Delta_nk][eV] (iverbosity=2)
4 prefix.acon_aniso_gap0_XX ! Re[\Delta_nk(0)][eV] distribution
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```

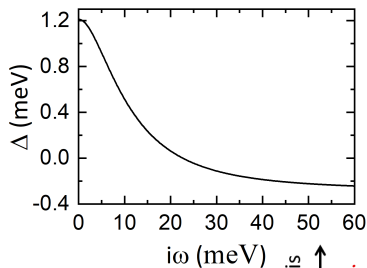
# Analytic Continuation to Real Axis



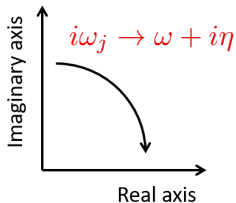
Isotropic case in Pb

Figures from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

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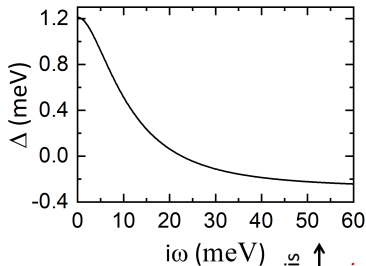
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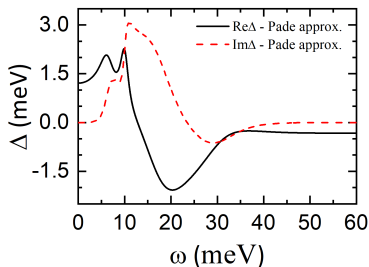
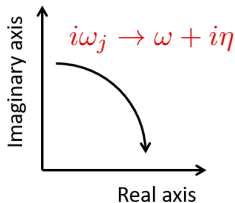
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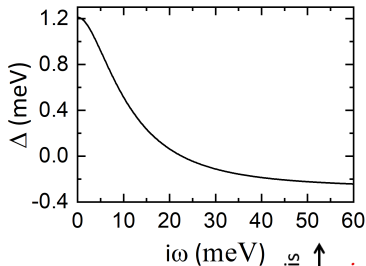


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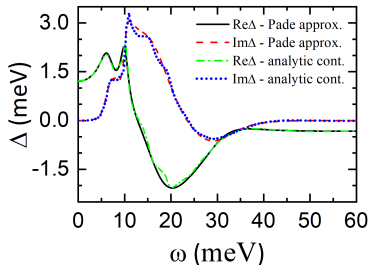
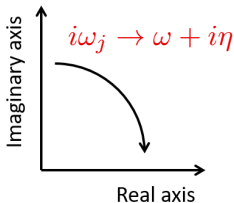


Figures from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

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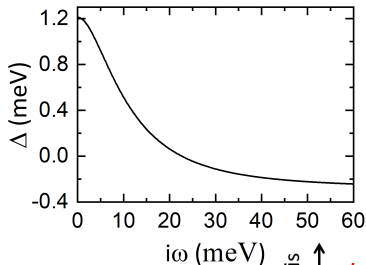


Isotropic case in Pb



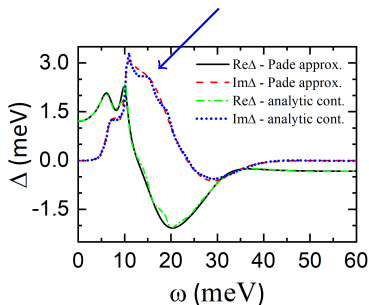
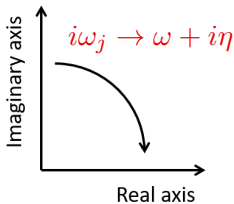
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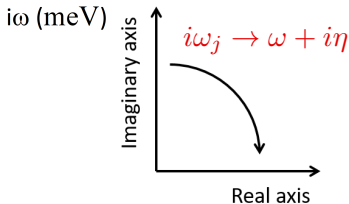
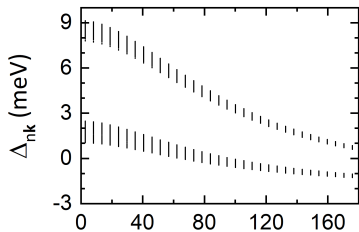
Isotropic case in Pb

structure in the real axis solutions on the scale of the phonon energy



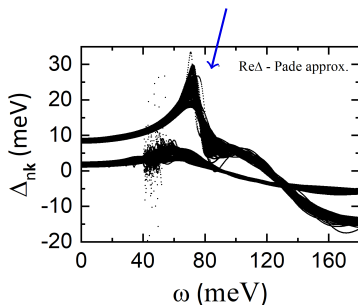
Figures from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

# Analytic Continuation to Real Axis in MgB<sub>2</sub>



Anisotropic case in MgB<sub>2</sub>

structure in the real axis solutions on the scale of the phonon energy



Figures from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

# Excitation Spectrum of a Superconductor

- The single-particle Green's function on real axis is given by:

$$\hat{G}_{n\mathbf{k}}(\omega) = \frac{\omega Z_{n\mathbf{k}}(\omega)\hat{\tau}_0 + (\epsilon_{n\mathbf{k}} - \epsilon_F)\hat{\tau}_3 + \Delta_{n\mathbf{k}}(\omega)Z_{n\mathbf{k}}(\omega)\hat{\tau}_1}{[\omega Z_{n\mathbf{k}}(\omega)]^2 - (\epsilon_{n\mathbf{k}} - \epsilon_F)^2 - [Z_{n\mathbf{k}}(\omega)\Delta_{n\mathbf{k}}(\omega)]^2}$$

$$\hat{\tau}_0 = \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix} \quad \hat{\tau}_1 = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix} \quad \hat{\tau}_3 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}$$

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- The pole positions are:  $E_{n\mathbf{k}}^2 = \frac{(\epsilon_{n\mathbf{k}} - \epsilon_F)^2}{Z_{n\mathbf{k}}^2(E_{n\mathbf{k}})} + \Delta_{n\mathbf{k}}^2(E_{n\mathbf{k}})$



# Superconducting Quasiparticle Energy and Lifetime

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This identity defines the leading edge  $\Delta_{n\mathbf{k}}$  of the superconducting gap

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binding energy for electrons  
in a Cooper pair

# Superconducting Quasiparticle Density of States

- The superconducting quasiparticle density of states can be deduced from:

$$\frac{N_{n\mathbf{k},S}(\omega)}{N_F} = -\frac{1}{\pi} \int_{-\infty}^{\infty} d\epsilon_{n\mathbf{k}} \text{Im} G_{n\mathbf{k}}^{11}(\omega)$$

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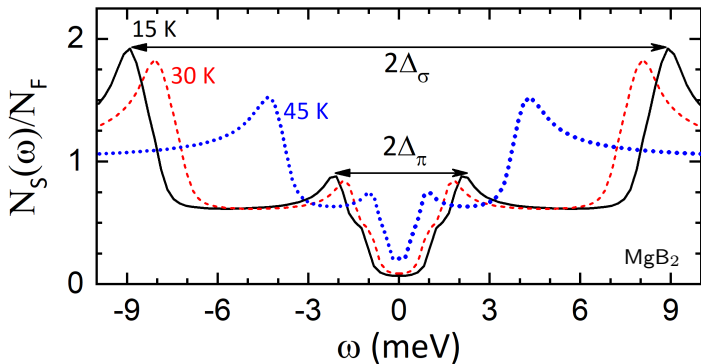
$$\frac{N_{n\mathbf{k},S}(\omega)}{N_F} = \text{Re} \left[ \omega / \sqrt{\omega^2 - \Delta_{n\mathbf{k}}^2(\omega)} \right]$$

- Averaging over the Fermi surface leads to:

$$\frac{N_S(\omega)}{N_F} = \sum_n \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \frac{\delta(\epsilon_{n\mathbf{k}} - \epsilon_F)}{N_F} \text{Re} \left[ \omega / \sqrt{\omega^2 - \Delta_{n\mathbf{k}}^2(\omega)} \right]$$



# Superconducting Quasiparticle Density of States



Figures from Margine and Giustino, Phys. Rev. B 87, 024505 (2013)

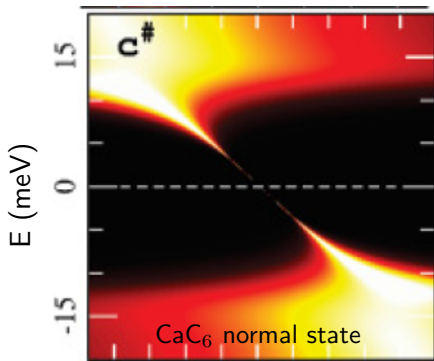
# Spectral Function

$$A_{n\mathbf{k}}(\omega) = -\frac{1}{\pi} \text{Im}G_{n\mathbf{k}}^{11}(\omega)$$

Figures from Sanna et al, Phys. Rev. B 85, 184514 (2012)

# Spectral Function

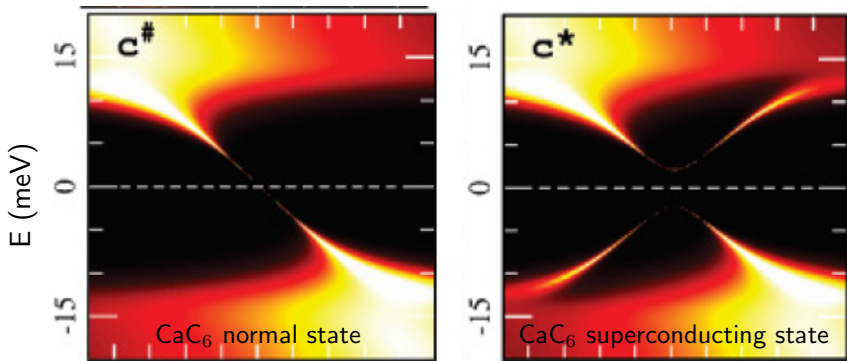
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# Miscellaneous

- `wscut` is normally set to 4 to 10 times largest phonon energy
- `wscut` is ignored if the frequencies on the imaginary axis are given with `nswi`
- `laniso/liso` requires `eliashberg`
- `lpade` requires `limag`
- `lacon` requires `limag` and `lpade`
- $T_c$  evaluated with Allen-Dynes formula can be used as a guide for defining the temperatures at which to evaluate the Migdal-Eliashberg eqs.
- `ephwrite` requires uniform fine  $\mathbf{k}$  or  $\mathbf{q}$  grids and `nkf1,nkf2,nkf3` to be multiple of `nqf1,nqf2,nqf3`
- `.ephmatXX`, `.egnv`, `.freq`, and `.ikmap` files need to be generated whenever  $\mathbf{k}$  or  $\mathbf{q}$  fine grid is changed

# Miscellaneous

- `imag_read` requires `limag` and `laniso`
- `imag_read` allows the code to read from file the superconducting gap and renormalization function on the imaginary axis at specific temperature `XX` from file `.imag_aniso_XX`. The temperature is specified as `tempsmin = XX` or `temps(1) = XX`.
- `imag_read` can be used to:
  - (1) solve the anisotropic Migdal-Eliashberg equations on the imaginary axis at temperatures greater than `XX` using as a starting point the superconducting gap estimated at temperature `XX`.
  - (2) obtain the solutions of the Migdal-Eliashberg equations on the real axis with `lpade` or `lacon` starting from the imaginary axis solutions at temperature `XX`;
  - (3) write to file the superconducting gap on the Fermi surface in cube format at temperature `XX` for `iverbosity = 2`.

# References

- E. R. Margine and F. Giustino, Phys. Rev. B 87, 024505 (2013) [\[link\]](#)
- S. Poncé, E. R. Margine, C. Verdi, and F. Giustino, Comput. Phys. Commun. 209, 116 (2016) [\[link\]](#)
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- P. B. Allen, and B. Mitrović, Solid State Phys. 37, 1 (1982) [\[link\]](#)
- C. R. Leavens and D. S. Ritchie, Solid State Commun. 53, 137 (1985) [\[link\]](#)
- F. Marsiglio, M. Schossmann, and J. P. Carbotte, Phys. Rev. B 37, 4965 (1988) [\[link\]](#)

## More info



<http://epw.org.uk>



<http://epwforum.uk>



<https://gitlab.com/QEF/q-e>